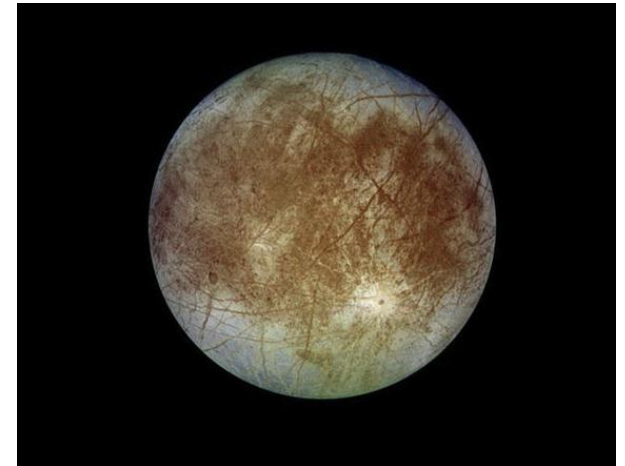


Magnesium Diboride (MgB₂)-Based Bolometer Array for Far IR Thermal Imaging and Fourier Transform Spectroscopy Applications

Rationale:

- Need for more sensitive detectors:
 - Thermal mapping of cold planetary bodies
 - Study of atmospheric molecular species using Fourier Transform Spectrometer (FTS) applications.
 - Outer planets and icy satellites missions
 - Future Uranus, Neptune, Titan missions

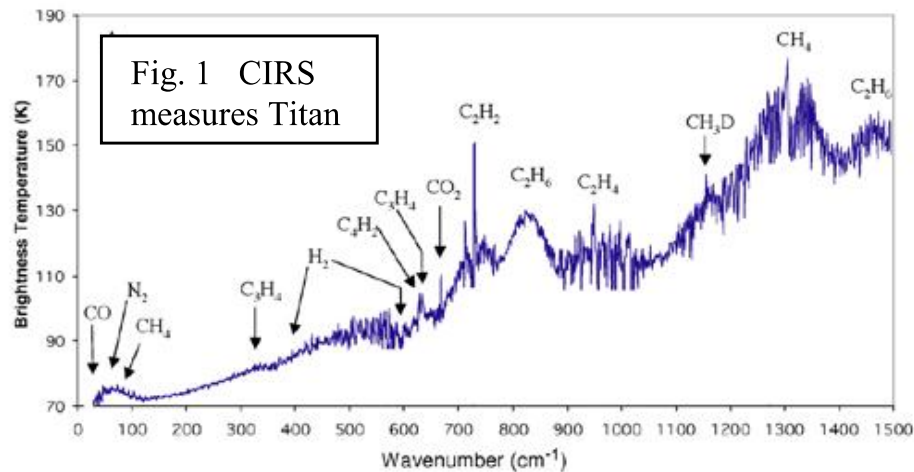


Europa

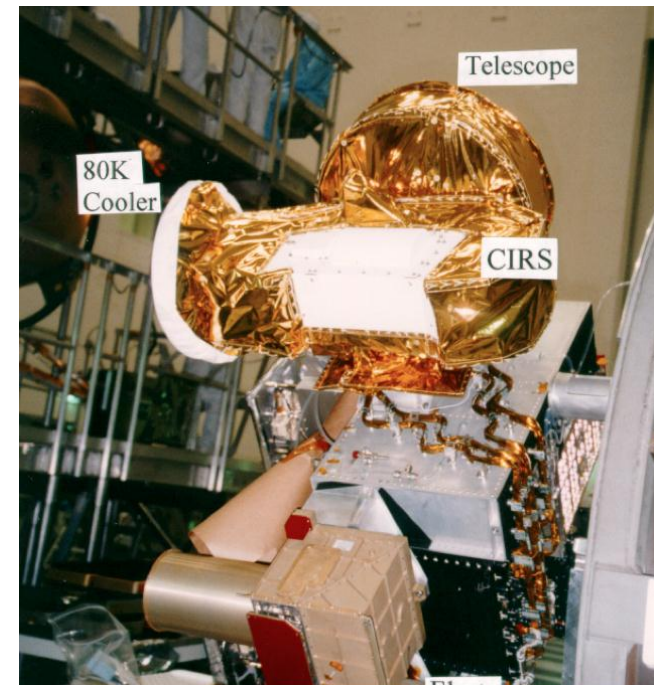


Enceladus

FTS: Improvement in the long wavelength portion of the spectrum



- Single pixels ---FTS
- Linear array ---Push broom configuration/
thermal mapping
- 2-D array--- Far-IR camera



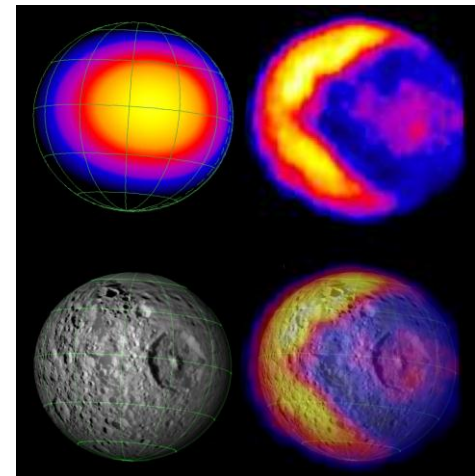
CIRS/Cassini instrument

What is available for far-IR investigation

- CIRS/Cassini:
 - FIR focal plane
 - Two single BiTe thermal detectors (thermopiles) operating at 170 K;
 - Each detector with a specific detectivity, D^* $\sim 3 \times 10^9 \text{ cmHz}^{1/2}/\text{W}$ near the low frequency end of a 0.4-to-30 Hz band pass.
- Commercially available pyroelectrics have a $D^* \sim 7 \times 10^8 \text{ cmHz}^{1/2}/\text{W}$ and operate at 300K
- YBCuO has good sensitivity $\sim 1 \times 10^{10} \text{ cmHz}^{1/2}/\text{W}$
 - But grows only on R-plane sapphire, LaAlO₃..
 - Difficult to architecture
 - Time constant $\geq 100 \text{ ms}$



Enceladus

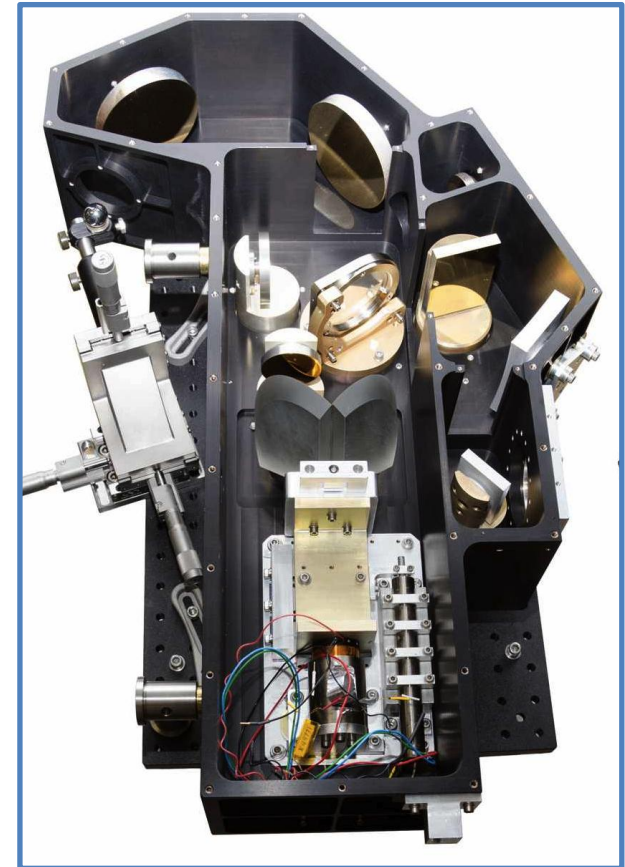


Mimas

- New smaller FTS under development (J. Brasunas is PI)
- Less than 1/2 the mass of CIRS/Cassini
- Baselines high Tc superconducting thermal detector at far-IR focal plane
 - Magnesium diboride (MgB₂)
 - Yttrium Barium Copper Oxide (YBCuO)

Parameter	CIRS	CIRS-lite
band-pass (μm)	7 to 1000	7 to 333
resolution (cm ⁻¹) apodized	0.5	0.125
telescope diameter (cm)	50	15
detectors	HgCdTe thermopile	HgCdTe high Tc
detector temperature (K)	75 and 170	75 and 89
optics temperature (K)	170	~150
point-able mirror	no	TBD 1 kg
footprint (km @ 250 km)	1 & 0.05	1 & 0.4
mass (kg)	43	15 to 20

Brasunas et al.

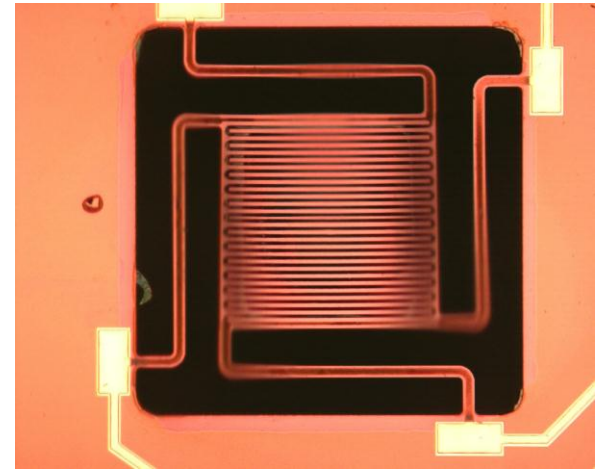


Cirs-Lite

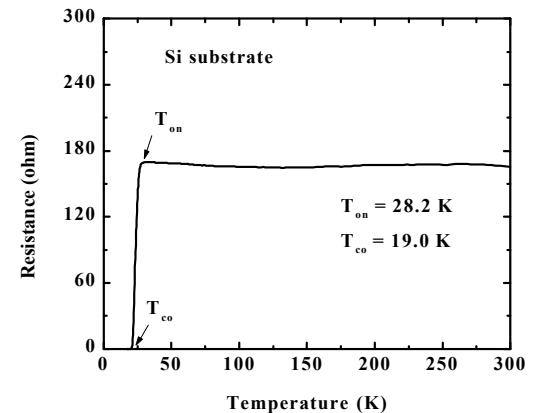
MgB2 far-IR detector development

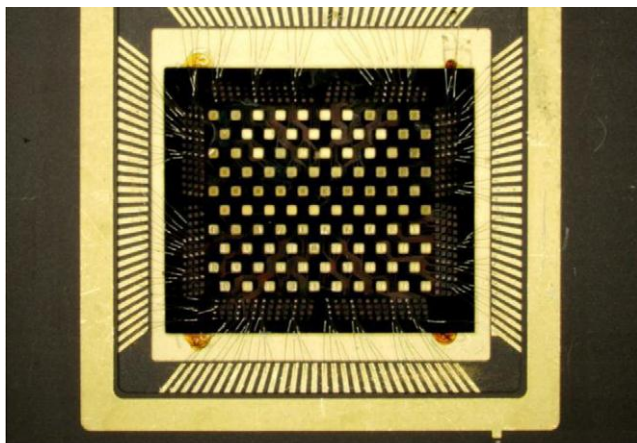
Our interest in MgB2:

- Simple binary, intermetallic compound as opposed to YBCuO
- Superconducts near 38 K
- Grows on Si- SiN substrates and not exclusively on R-plane Al₂O₃ or LaAlO₃. Easy to architecture (R and bias).
- Has the desirable properties for the development of sensitive high-T_c bolometers
- Recent positive advances in deposition techniques
- Quality polycrystalline MgB₂ with T_c near 38K



MgB2 on SiN pixel





(b)

Fabricated 10x10 MgB2 array (250 x 250 μm each)

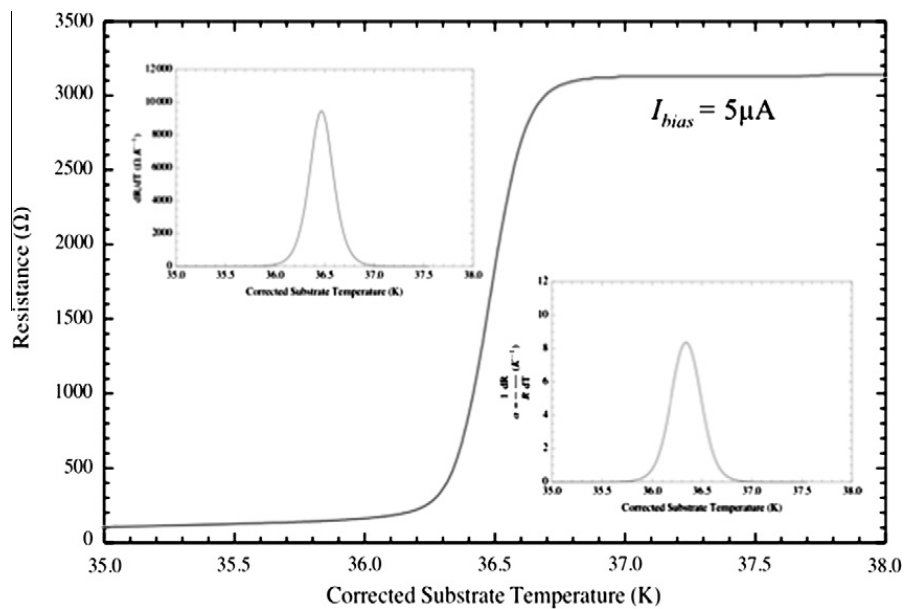
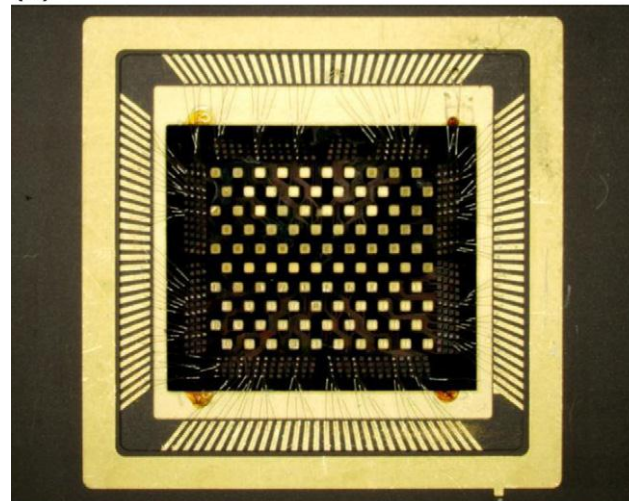


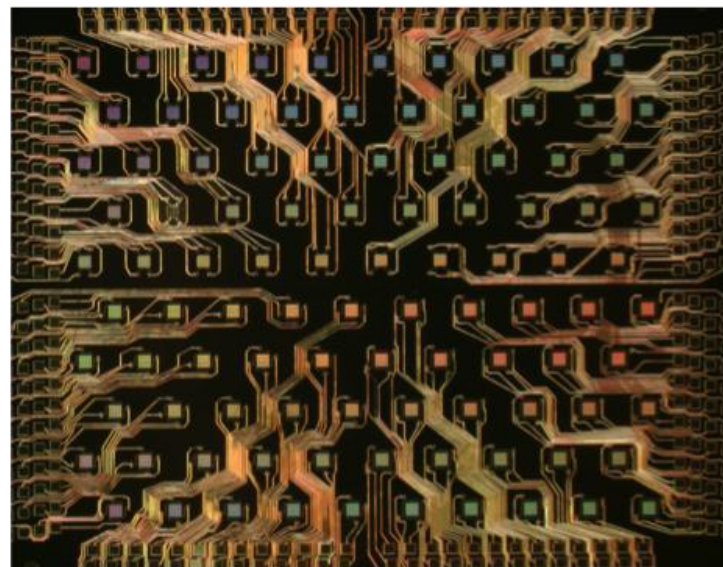
Fig. 3. R , dR/dT and α as a function of corrected substrate temperature.

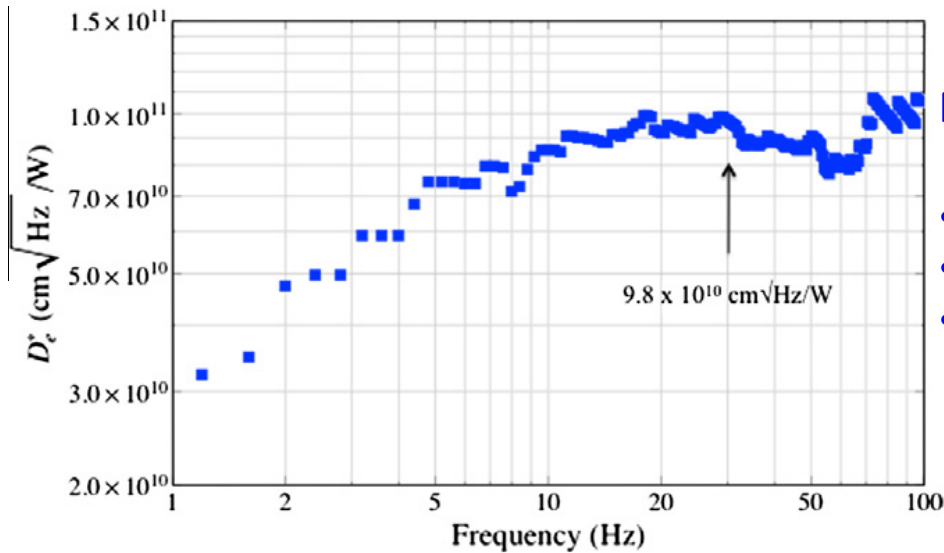
(a)



(b)

Thin film MgB2 bolometer with resistive meander line (no absorber)





MgB2:

- Better sensitivity than thermopiles and YBCuO
- Better time constant
- Easier to produce single and 2-D arrays

Fig. 7. Electrical detectivity, D_e^* , as a function of frequency.

More at:

Physica C 483 (2012) 119–126

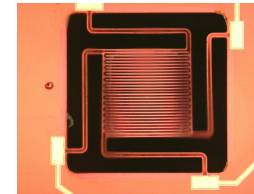
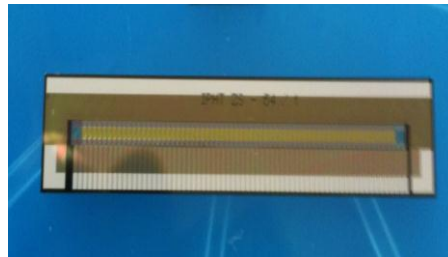
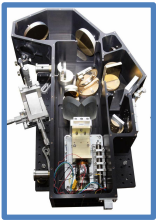
B. Lakew, S. Aslam *et al.*

	Temp K	Optical Sensitivity D^*
Thermopile	150-170	3×10^9
YBCuO	90	1×10^{10}
MgB2	38	$\sim 8.3 \times 10^{10}$

In conclusion:

MgB₂ is a better alternative to thermopiles and YBCuO
Easier to produce using MEMS processing

Challenge: Needs a cryocooler to operate at the transition.



TRL3 + cryocooler